

Unitarization and Causalization of Non-local quantum field theories by Classicalization

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Abstract

We suggest that classicalization can cure non-local quantum field theories from acausal divergences in scattering amplitudes, restoring unitarity and causality. In particular, in "trans-non-local" limit, the formation of non-perturbative classical configurations, called *classicalons*, in scatterings like $\phi\phi \rightarrow \phi\phi$, can avoid typical acausal divergences.

1 Introduction

Classicalization is an alternative "road" to an UV completion of quantum field theories with respect to Wilsonian one. Such a phenomena was conjectured by Dvali and collaborators, and a lot of examples seem to sustain this argument [1]. In this paper, we would like to argue possible connections among classicalization and non-local quantum field theories. To formulate a consistent quantum field theory without the locality principle is an old problem: it is well known as an "insidious problem". In fact, even if one can formulate a consistent model at tree level, unitarity and causality will be inevitably lost at quantum level [4, 5, 6, 7, 8, 9, 10]. For example, in a non-local scalar field theory, acausalities will inevitably appear in scatterings like $\phi\phi \rightarrow \phi\phi$ or $\phi\phi \rightarrow \phi\phi\phi\phi$. However, classicalization can offer a natural way-out to acausal divergences: the formation of classical extended objects of radius $R > \Lambda_{NL}^{-1}$ in scatterings with $E_{CM} > \Lambda_{NL}$, named classicalons, naturally avoids these infinities (where Λ_{NL} is the Non-locality effective scale). In particular, we will quantitatively focus on a particularly promising class of Non-local models studied by Eliezer, Woodard, Moffat, Kleppe, Evens and Joglekar [4, 5, 6, 7, 8, 9, 10]².

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²An alternative approach to non-local QFT and non-local quantum gravity is considered in [11, 12, 13, 14, 15, 16, 17, 18, 19, 20]. The main difference with respect to our approach is the following: a infinite number of local gauge transformations are considered in their case rather than a non-linear one. I am grateful to Leonardo Modesto for discussions on these aspects.

2 EWMKEJ's model: nonlocal scalar field theory

Let us briefly review the EWMKEJ (*i.e.*, Eliezer, Woodard, Moffat, Kleppe, Evens, Joglekar) model of a nonlocal $\lambda\phi^4$ theory. We start from an action

$$S[\phi] = \mathcal{F}[\phi] + \mathcal{I}[\phi], \quad (1)$$

where $\mathcal{F}[\phi]$ is the free part, $\mathcal{I}[\phi]$ is the interaction part. We suppose analytic functional around the vacuum. \mathcal{F} has a general form

$$\mathcal{F}[\phi] = -\frac{1}{2} \int d^4x \phi_i F_{ij} \phi_j. \quad (2)$$

The action S is "nonlocalized" through a "smearing operator" \mathcal{E} . The EWMKEJ choice corresponds to an exponential smearing operator:

$$\mathcal{E} = \exp \left[\frac{\mathbf{F}}{2\Lambda_{\text{NL}}^2} \right], \quad (3)$$

Λ_{NL} is an effective scale of non-locality. Then, ϕ are smeared as

$$\hat{\phi}_i = \mathcal{E}_{ij}^{-1} \phi_j. \quad (4)$$

Let us define the operator

$$\mathcal{K} \equiv (\mathcal{E}^2 - I)F^{-1}. \quad (5)$$

Now, we introduce an auxiliary field φ_i for each matter field ϕ_i :

$$\hat{S}[\phi, \varphi] = \mathcal{F}[\hat{\phi}] - \mathcal{A}[\varphi] + \mathcal{I}[\phi + \varphi], \quad (6)$$

$$\mathcal{A}[\varphi] = -\frac{1}{2} \int d^4x \varphi_i \mathcal{K}_{ij}^{-1} \varphi_j. \quad (7)$$

The classical auxiliar field equation is

$$\frac{\delta \hat{S}[\phi, \varphi]}{\delta \varphi(x)} = 0. \quad (8)$$

The final nonlocal action is non-linearly obtained by substituting the solution of Eq. (8), into Eq. (6); *i.e* by substituting $\varphi_i = \mathcal{K}_{ij} \frac{\delta \mathcal{I}[\phi + \varphi]}{\delta \varphi_j}$.

2.1 Gauge symmetries

How can be constructed a local gauge symmetry in a non-local model? Gauge symmetry can be encoded in a nonlocal theory with a new nonlinear transformation rule. In fact, as shown in Ref.[7] for the scalar theory, if an infinitesimal transformation $\delta\phi_i = T_i[\phi]$

generates a symmetry of the local action $S[\phi]$, then a transformation $\hat{\delta}\phi_i = \mathcal{E}_{ij}^2 T_j[\phi + \varphi[\phi]]$ generates a symmetry for the corresponding nonlocal action $\hat{S}[\phi]$. In a broad sense, the procedure for obtaining a nonlocal theory preserves a deformed version of the usual continuous symmetry, and we can write

$$\hat{\delta}\varphi_i[\phi] = (I - \mathcal{E}^2)_{ij} T_j[\phi + \varphi[\phi]] - K_{ij}[\phi + \varphi[\phi]] \frac{\delta T_k}{\delta\phi_j}[\phi + \varphi[\phi]] \mathcal{E}_{kl}^2 \frac{\delta \hat{S}[\phi]}{\delta\phi_l}, \quad (9)$$

$$K_{ij}^{-1}[\phi] = \mathcal{K}_{ij}^{-1} - \frac{\delta^2 \mathcal{I}[\phi]}{\delta\phi_i \delta\phi_j}. \quad (10)$$

2.2 Quantization in the EWMKEJ model

Consider the vacuum expectation value of an arbitrary operator \mathcal{O} :

$$\langle \mathcal{T}(\mathcal{O}[\phi]) \rangle_{\mathcal{E}} = \int \mathcal{D}\phi m[\phi] (GF) \mathcal{O}[\hat{\phi}] e^{i\hat{S}[\phi]} \quad (11)$$

(\mathcal{T} is the time-ordering operator, and GF is the gauge fixing). In this definition, \mathcal{O} is nonlocally regulated and eq.(11) defines the quantization in non-local model.

A consistent quantization of EWMKEJ requests the existence of the measure factor $m[\phi]$ and the gauge fixing. A measure functional is necessary in order to preserve unitarity. Unitarity of a nonlocal quantum field theory was discussed in papers cited above. In particular, a large subspace \mathcal{M} in the Fock space posses unitarity at three level. As in local QFT, \mathcal{M} can also have unphysical polarizations as BRST ghost fields. On the other hand, non-linear gauge invariance guarantees BRST ghosts' decoupling on shells.

So, the EWMKEJ procedure starts from a local QFT in order to obtain a non-local deformation. Generically, the starting QFT has continuos transformations $\delta\phi_i = T_i[\phi]$ of the local action $S[\phi]$. EWMKEJ procedure generates corresponding transformations $\hat{\delta}\phi_i = \mathcal{E}_{ij}^2 T_j[\phi + \varphi[\phi]]$ for the non-local QFT. However, this transformation has to preserve $\mathcal{D}\phi m[\phi]$, i.e $\hat{\delta}[\mathcal{D}\phi m[\phi]] = 0$. Such a condition corresponds to

$$\hat{\delta}[\log(m[\phi])] = -\text{Tr} \left[\frac{\delta \hat{\delta}\phi_i}{\delta\phi_j} \right] = -\text{Tr} \left[\mathcal{E}_{ik}^2 \frac{\delta T_k}{\delta\phi_l}[\phi + \varphi[\phi]] K_{lk}[\phi + \varphi[\phi]] \mathcal{K}_{kj}^{-1} \right]. \quad (12)$$

Under this quantization procedure, we can recover Feynman rules of $\hat{S}[\phi, \varphi]$ as simple extension of usual ones: propagators are smeared by a factor \mathcal{E}^2 , as mentioned above. The φ are auxiliary fields propagating only off-shell, because they are projected-out by solutions of classical field equations $\varphi[\phi]$.

2.3 Non-local Feynman rules

The "funny trick" of the auxiliar field allows to obtain simple Feynman rules, as understood deformations of usual ones. Let us resume the new prescriptions demonstrated in papers cited above:

- i) the vertices are unchanged;
- ii) the smeared propagator for the fields ϕ reads as

$$-\frac{i\mathcal{E}^2}{(F + i\epsilon)} \quad (13)$$

- iii) The smeared propagators for the auxiliary fields φ are

$$-\frac{i[I - \mathcal{E}^2]}{(F + i\epsilon)} \quad (14)$$

where I is the identity-operator.

Let us choice

$$F = \square + m^2 \quad (15)$$

We can conveniently write Feynman rules in momentum space are as follows.

- ii)

$$i \frac{\exp\left(\frac{-(p^2 - m^2)}{\Lambda_{NL}^2}\right)}{(p^2 - m^2 + i\epsilon)} \quad (16)$$

- iii)

$$i \frac{\left[I - \exp\left(\frac{-(p^2 - m^2)}{\Lambda_{NL}^2}\right) \right]}{(p^2 - m^2 + i\epsilon)} \quad (17)$$

2.4 Bogoliubov-Shirkov causality conditions

In order to test causality and unitarity at all the orders of the perturbation series, one can find recursive relations on the S-matrices. Such conditions were discussed by Bogoliubov and Shirkov in their book on QFT [21].

2.4.1 Unitarity and Causality: definitions

It is useful to remind what unitarity and causality impose on the S-matrix.

Unitarity means that the total probability of processes equal to one: S -matrix has to satisfy the condition

$$S S^\dagger = S^\dagger S = I. \quad (18)$$

Causality can be reformulated as the cluster decomposition principle on the S-matrix [22]. If multi-particle transitions $A_1 \rightarrow B_1, A_2 \rightarrow B_2, \dots, A_n \rightarrow B_n$ are studied in N different laboratories, with different positions $z_{1,\dots,n}$ in the space-time with $(z_i - z_j)^2 < 0$ ($i \neq j, i, j = 1, \dots, n$), then the S-matrix will be decomposed as

$$\mathcal{S}_{B_1+B_2+\dots+B_n, A_1+A_2+\dots+A_n} = \mathcal{S}_{B_1 A_1} \mathcal{S}_{B_2 A_2} \dots \mathcal{S}_{B_n A_n}. \quad (19)$$

Rel.(19), is strictly connected to the hypothesis that quantum fields commute as

$$[\phi(z_i), \phi(z_j)] = 0, \quad (z_i - z_j)^2 < 0, \quad (20)$$

which is equivalent to a microcausality condition on the S-matrix:

$$\frac{\delta}{\delta\phi(z_i)} \left(\frac{\delta\mathcal{S}[\phi]}{\delta\phi(z_j)} \mathcal{S}^\dagger[\phi] \right) = 0, \quad \text{for } z_i < z_j. \quad (21)$$

2.4.2 Recursive relations as a "Test-Bed" for unitarity and causality

Let us perform a Dyson expansion of the S-matrix with respect to the couplings $c(z)$ promoted to spacetime fields [21]:

$$\mathcal{S}[c(x)] = I + \sum_{n=1}^{\infty} \frac{1}{n!} \int dz_1 \dots dz_n \mathcal{T}\{\mathcal{S}_n(z_1, \dots, z_n) c(z_1) \dots c(z_n)\}. \quad (22)$$

Let us note that conditions (18) and (21) can be rewritten in terms of $c(z)$, through a functional Legendre transform $\phi(z) \rightarrow c(z)$. Then, for $c(z) \rightarrow c = \text{const}$, we can insert the expansion (22) into (18) and (21), reverting the usual expansion. These allow to obtain the following recursive relations, for perturbation theory:

$$\mathcal{R}_n = i\mathcal{S}_{n+1}(y, z_1, \dots, z_n) + i \sum_{0 \leq k \leq n-1} \mathcal{P}\{\mathcal{S}_{k+1}(y, z_1, \dots, z_k) \mathcal{S}_{n-k}^\dagger(z_{k+1}, \dots, z_n)\}, \quad (23)$$

and

$$\mathcal{S}_n(z_1, \dots, z_n) + \mathcal{S}_n^\dagger(z_1, \dots, z_n) + \sum_{1 \leq k \leq n-1} \mathcal{P}\{\mathcal{S}_k(z_1, \dots, z_k) \mathcal{S}_{n-k}^\dagger(z_{k+1}, \dots, z_n)\} = 0, \quad (24)$$

where $\mathcal{P}\{\}$ is the sum over all partitions of $\{z_1, \dots, z_n\}$ into k and $n - k$ elements. For example, $\{z_1, \dots, z_k\}, \{z_{k+1}, \dots, z_n\}$ and so on.

The causality conditions for the first two orders of the perturbation theory are

$$\mathcal{R}_1(x, y) = i \left[\mathcal{S}_2(x, y) + \mathcal{S}_1(x) \mathcal{S}_1^\dagger(y) \right] = 0, \quad (25)$$

$$\mathcal{R}_2(x, y) = i \left[\mathcal{S}_3(x, y, z) + \mathcal{S}_1(x) \mathcal{S}_2^\dagger(y, z) + \mathcal{S}_2(x, y) \mathcal{S}_1^\dagger(z) + \mathcal{S}_2(x, z) \mathcal{S}_1^\dagger(y) \right] = 0. \quad (26)$$

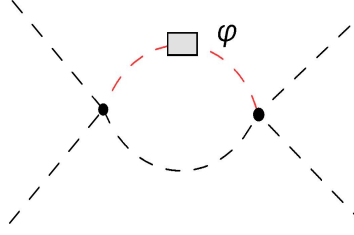


Figure 1: Example of 1-loop contribution to $\phi\phi \rightarrow \phi\phi$ scattering leading to acausal divergences. In this diagram, one off-shell auxiliary field φ (red dashed lines with a box) and an ordinary one are inside the one-loop. This diagram is related to the interaction vertex $\lambda\phi^3\varphi$ inside $\mathcal{I}[\phi + \varphi]$.

On the other hand, unitarity is expressed at the first order as

$$\mathcal{S}_1(x) + \mathcal{S}_1^\dagger(x) = 0, \quad (27)$$

$$\mathcal{S}_2(x, y) + \mathcal{S}_2^\dagger(x, y) + \mathcal{S}_1(x)\mathcal{S}_1^\dagger(y) + \mathcal{S}_1(y)\mathcal{S}_1^\dagger(x) = 0. \quad (28)$$

An alternative useful way to rewrite (25)-(26) is

$$\mathcal{R}_1 = \int d^4x d^4y [\theta(x_0 - y_0)\mathcal{R}_1(x, y) + \theta(y_0 - x_0)\mathcal{R}_1(y, x)] = 0, \quad (29)$$

$$\mathcal{R}_2 = \int d^4x d^4y d^4z \mathcal{R}_2(x, y, z)\theta(x_0 - y_0)\theta(y_0 - z_0) + 5 \text{ symmetric terms} = 0. \quad (30)$$

Let us remark that these bounds are surely a "test-bed" for causality and unitarity violations. In fact, suppose to calculate a 1 and 2 vertices' amplitudes: from the *momentum dependence* of $\mathcal{S}_{1,2}$, implying a certain function of momentum in \mathcal{R}_1 , *one can manifestly see signals of causality or unitarity violations*. In fact, if the net contribution of the relevant amplitudes to \mathcal{R}_1 is proportional to a polynomial function of the Mandelstam variables $\mathcal{X}^n/\Lambda_{NL}^{2n}$ with $n > 1$ ($\mathcal{X} = s, t, u$), for $\mathcal{X} \gg \Lambda_{NL}^2$ a breakdown of causality will occur. In order to have a causal theory, \mathcal{R}_1 has to be zero, as well as $\mathcal{R}_{2,\dots,n}$. Analogous for unitarity. This last condition is violated in non-local models like EWMKEJ, as we will see in the next sections. In fact, the presence of off-shell auxiliary fields will introduce extra un-balanced contributions violating causality and unitarity bounds.

3 Acausal divergences in Scatterings and Classicalization

3.1 One-loop acausal diagrams in $\phi\phi \rightarrow \phi\phi$ scatterings

In this section, we show an explicit example of how unitarity and causality are lost at quantum level. In particular, we briefly review the case of a scattering $\phi\phi \rightarrow \phi\phi$

corrected by a loop of an auxiliary field φ and one field ϕ . This example was discussed in Ref.[9, 10], and Ref.[23] in a susy generalization.

By assuming the massless case $m = 0$, the (renormalized) amplitude is

$$\mathcal{A}(s, t, u) = \frac{9\lambda^2}{4\pi^2} \sum_{\mathcal{X}=s,t,u} \sum_{n=0}^{\infty} a_n \left(\frac{\mathcal{X}}{\Lambda_{NL}^2} \right)^n \quad (31)$$

where

$$a_n = \frac{1}{2^n n(n+1)!} (2^n - 1)$$

and $\mathcal{X} = s, t, u$ are the Mandelstam variables.

Let us comment that the reintroduction of the mass parameter will complicate the form of (31), but essentially this is not important for our purposes. Anyway, one can handle also masses using Schwinger parameters, as usually done for local QFT. One can rewrite the amplitudes as the following integrals in s -channel:

$$\mathcal{A}(s) = \frac{9\lambda^2}{4\pi^2} \int_0^{1/2} dx \int_{\frac{1}{(1-x)}}^{\frac{1}{x}} \frac{d\zeta}{\zeta} \exp \left\{ -\frac{\zeta}{\Lambda_{NL}^2} (m^2 - x(1-x)s) \right\}. \quad (32)$$

The complete expansion is complicated, but we can perform an asymptotic expansion around $s = 0$, as

$$\mathcal{A}(s) \sim \sum_{n=0}^2 a_n(m, \Lambda_{NL}) s^n + O(s^3), \quad (33)$$

in which the coefficients $a_{0,1,2}$ are expressed as the following integrals

$$a_0(m, \Lambda_{NL}) = \frac{9\lambda^2}{4\pi^2} \int_0^{1/2} dx \int_{\frac{1}{(1-x)}}^{\frac{1}{x}} \frac{d\zeta}{\zeta} e^{-\frac{m^2 \zeta}{\Lambda_{NL}^2}}, \quad (34)$$

$$a_1(m, \Lambda_{NL}) = \frac{9\lambda^2}{4\pi^2} \int_0^{1/2} dx \int_{\frac{1}{(1-x)}}^{\frac{1}{x}} d\zeta \frac{x(1-x)}{\Lambda_{NL}^2} e^{-\frac{m^2 \zeta}{\Lambda_{NL}^2}}, \quad (35)$$

$$a_2(m, \Lambda_{NL}) = \frac{9\lambda^2}{4\pi^2} \int_0^{1/2} dx \int_{\frac{1}{(1-x)}}^{\frac{1}{x}} \zeta d\zeta \frac{x^2(1-x)^2}{2\Lambda_{NL}^4} e^{-\frac{m^2 \zeta}{\Lambda_{NL}^2}}. \quad (36)$$

The zeroth and first order of the expansion (33) can be cancelled in the renormalization procedure, because they are just constants.

Let us note that new polynomial terms "strongly" violate unitarity and causality relations in the "trans-nonlocal regime" $E \gg \Lambda_{NL}$ (or $\mathcal{X} \gg \Lambda_{NL}^2$). In fact, such a scattering contributes only to $\int dz_1 dz_2 \mathcal{S}_2(z_1, z_2)$ and not to $\int dz_1 dz_2 \mathcal{T}\{\mathcal{S}_1(z_1) \mathcal{S}_1^\dagger(z_2)\}$: the net $\mathcal{R}_1(\mathcal{X})$ is momentum-dependent as $\sim \mathcal{X}^2/\Lambda_{NL}^4 + O(\mathcal{X}^3/\Lambda_{NL}^6)$. This contribution is not balanced by other correspondent ones. However, this conclusion is right in the

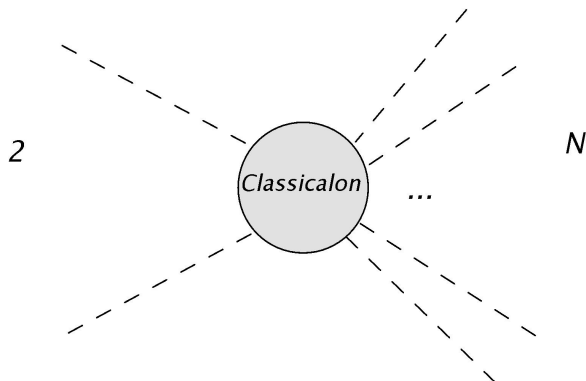


Figure 2: Classicalon' production in a scattering of two particles.

Wilsonian UV completion, but not for a UV Classicalization. In fact, even if $E > \Lambda_{NL}$, the formation of a non-perturbative extended classic object called "classicalon" can cutoff the minimal length probed by the scattering to the size R of the classicalon. If this size $R > \Lambda_{NL}^{-1}$, unitarity and causality are not lost in these channels.

3.2 Classicalization

The qualitative idea of classicalization is the following, in the limits of $\sqrt{s} > \Lambda_{critical} = l_{critical}^{-1}$ (as well as limits in t- and u- channels), (where $\Lambda_{critical}, l_{critical}$ are a critical energy and length scales respectively) the production of an extended classical configurations, called "classicalon" starts dominating the high energy amplitudes. For example in a $2 \rightarrow 2$ process, energy-divergences are exponentially suppressed: new channels (shown in Fig.2)

$$2 \rightarrow \text{Classicalon} \rightarrow N$$

start to dominate for $\sqrt{s} > \Lambda_{critical}$. As a consequence, in the limit of $\sqrt{s} > \Lambda_{critical}$, the scattering cannot proceed down to scales $l < l_{critical}$, but it will sustain the creation of an extended object of size $R > l_{critical}$. As a consequence, a scattering with $\sqrt{s} \gg \Lambda_{critical}$ cannot probe distances $l \ll R$! Clearly, this will signify that if classical production dominates in our non-local QFT, length-scales down the non-locality length-scale will never be reached, *i.e* no-break down of unitarity and causality in scattering amplitudes!

Since we are talking about the formation of a classical configuration, one of the most appropriate languages, in order to describe classicalons, is the path integral formalism [2]. The idea is to study quantum fluctuations around classicalon rather than around

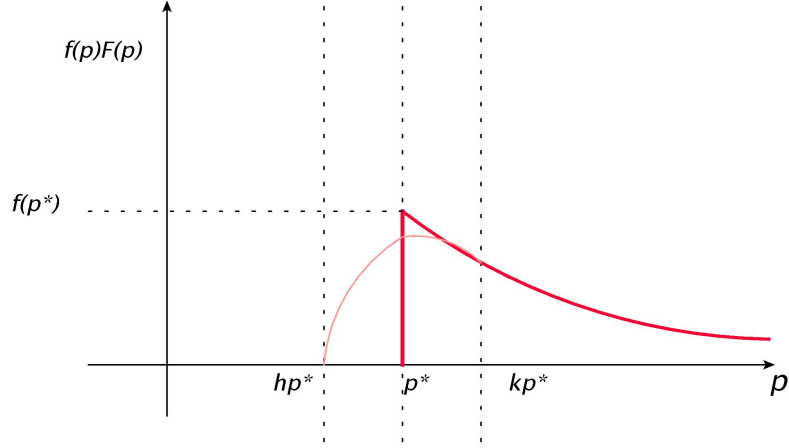


Figure 3: Qualitative examples of $f(p)F(p)$ distributions (where $F(p) = \mathcal{F}(p - p^*)$): in thick red the one with ansatz $\mathcal{F} = \Theta(p - p^*)$ while in light red and example of smoothed distributions converging to the Heaviside case for $p > kp^*$ and $p < hp^*$.

the trivial vacuum³. We will argument through path integral language how classicalons can unitarize and causalize scattering amplitudes in non-local models. Let us explicitly rewrite the lagrangian as

$$\mathcal{L} = -\frac{1}{2}e^{-\frac{1}{2}\frac{\square+m^2}{\Lambda_{NL}^2}}\phi(\square+m^2)e^{-\frac{1}{2}\frac{\square+m^2}{\Lambda_{NL}^2}}\phi - \frac{1}{2}\varphi(\phi)\frac{\square+m^2}{I - e^{\frac{\square+m^2}{\Lambda_{NL}^2}}}\varphi(\phi) - V(\phi + \varphi(\phi)) \quad (37)$$

and let us neglect the mass parameter, assumed $m \ll \Lambda_{NL}$ (this assumption will not be important for our arguments):

$$\mathcal{L} = -\frac{1}{2}e^{-\frac{1}{2}\frac{\square}{\Lambda_{NL}^2}}\phi(\square)e^{-\frac{1}{2}\frac{\square}{\Lambda_{NL}^2}}\phi - \frac{1}{2}\varphi(\phi)\frac{\square}{I - e^{\frac{\square}{\Lambda_{NL}^2}}}\varphi(\phi) - V(\phi + \varphi(\phi)) \quad (38)$$

where $\varphi(\phi)$ is a functional containing the derivatives of ϕ rather than polynomial terms. The associated generating function can be formally written as

$$Z[J] = \int \mathcal{D}\phi m[\phi] \exp\left(-\int d^4x (\mathcal{L} - J\phi)\right) \quad (39)$$

where J is the current, $m[\phi]$ the path integral measure. We expand around the classical solution ϕ_0 , $\phi = \phi_0 + \phi_1$. Integrating by part, we can rewrite inside the integral

$$\mathcal{L}(\phi) \rightarrow \mathcal{L}(\phi_0) - \phi_1 M \delta(x - y) + \frac{1}{2}\phi_1 H \phi_1 + \mathcal{L}_{int}(\phi_1)$$

where $M\delta(x-y)$ is the source in the Equation of Motion for ϕ , H is an infinite derivative operator. Another useful operation inside the integral is

$$J(x)\phi(x) \rightarrow (J(x) - M\delta(x - y))\phi_1(x)$$

³In a broad sense, classicalons can be considered as "brothers" of solitons, instantons and other non-perturbative solutions. Other intriguing implications of non-perturbative solutions called "exotic instantons" were recently studied in [51, 51, 52, 53, 54, 55, 56, 57, 58, 59, 60].

As a consequence, the path integral can be rewritten as

$$Z[J] \sim \int dR d^4y \mathcal{Y}(R, y) \frac{e^{-S[\phi_0(R, y)]}}{\sqrt{\text{Det}(H_{R, y})}} e^{-S_{int}[\frac{\delta}{\delta J}]} e^{\frac{1}{2} \int d^4x_1 d^4x_2 J(x_1) \Delta_{R, y}(x_1, x_2) J(x_2)} \quad (40)$$

where R, y are the coordinates parametrizing the classicalon solution, $\mathcal{Y}(R, y)$ is a measure correspondent to classicalon and the starting $m[\phi]$, $\Delta_{R, y}$ is the inverse function of $H_{R, y}$. Let us note that R parametrizes the radius of the classicalon.

All relevant informations about the classicalon are contained in the "Green problem" of the propagator $\Delta_{R, y}$:

$$H_{R, y} \Delta_{R, y}(\zeta, z) = \delta(\zeta - z) \quad (41)$$

Despite to the technical difficulties to find out an explicit expression for an infinite derivative operator, it is enough for our purposes to study the asymptotic limits of $H_{R, y}$. In particular, one can argue that for $r \rightarrow \infty$, we are OUT of interactions so that the effective lagrangian is asymptotically

$$\begin{aligned} \mathcal{L} &\rightarrow -\frac{1}{2} e^{-\frac{1}{2} \frac{\square}{\Lambda_{NL}^2}} \phi \square e^{-\frac{1}{2} \frac{\square}{\Lambda_{NL}^2}} \phi \\ &= -\left(1 - \frac{\square}{2\Lambda_{NL}^2} + \frac{\square^2}{4\Lambda_{NL}^4} + \dots + (-1)^n \frac{\square^n}{2^n \Lambda_{NL}^{2n}}\right) \phi \square \left(1 - \frac{\square}{2\Lambda_{NL}^2} + \frac{\square^2}{4\Lambda_{NL}^4} + \dots + (-1)^n \frac{\square^n}{2^n \Lambda_{NL}^{2n}}\right) \phi \end{aligned} \quad (42)$$

The corresponding Eulero-Lagrange equation can be rewritten as an expansion series

$$\sum_{m=0}^{\infty} A_m \square^{m+1} \phi = 0 \quad (43)$$

where the precise form of A_m is not reported because of not important for our following arguments ($A_0 = 1$). Clearly, in the "deep IR" limit the equation becomes just $\square \phi = 0$. In fact one can perform a Fourier transform of (48) so that

$$\lim_{p \rightarrow 0} \sum A_m p^{2m+2} \phi \sim p^2 \phi + O(p^4) \phi \quad (44)$$

This can be also easily understood as a limit $\square \rightarrow 0$ of (42): $e^{-\square/2\Lambda_{NL}^2} \rightarrow 1$, $\mathcal{L} \rightarrow -\frac{1}{2} \phi \square \phi$.

As a consequence, for $r \rightarrow \infty$, $H_{R, y}$ can be rewritten in spherical coordinates as

$$\begin{aligned} H_{R, y} &\rightarrow -\sum_{k=0}^{2n+1} \frac{(2n+1)!}{k!(2n+1-k)!} \frac{\partial^{4n+2-2k}}{\partial r^{4n+2-2k}} \sum_{j=0}^k \frac{(-1)^{k-j} 3^j k!}{j!(k-j)!} \frac{1}{r^j} \frac{\partial^j}{\partial r^j} \frac{L^{2(k-j)}}{r^{2(k-j)}} \\ &\quad -\sum_{k=0}^{2n+1} \frac{(2n+1)!}{k!(2n+1-k)!} \frac{\partial^{4n+2-2k}}{\partial r^{4n+2-2k}} \sum_{j=0}^k \frac{(-1)^{k-j} 3^{k-j} k!}{j!(k-j)!} \frac{L^{2j}}{r^{2j}} \frac{1}{r^{k-j}} \frac{\partial^{k-j}}{\partial r^{k-j}} \end{aligned} \quad (45)$$

$$\begin{aligned}
& - \sum_{k=0}^{2n+1} \frac{(2n+1)!}{k!(2n+1-k)!} \frac{L^{4n+2-2k}}{r^{4n+2-2k}} \sum_{j=0}^k \frac{(-1)^{k-j} 3^{k-j} k!}{j!(k-j)!} \frac{\partial^{2j}}{\partial r^{2j}} \frac{1}{r^{k-j}} \frac{\partial^{k-j}}{\partial r^{k-j}} \\
& - \sum_{k=0}^{2n+1} \frac{(2n+1)!}{k!(2n+1-k)!} \frac{L^{4n+2-2k}}{r^{4n+2-2k}} \sum_{j=0}^k \frac{(-1)^{k-j} 3^j k!}{j!(k-j)!} \frac{1}{r^j} \frac{\partial^j}{\partial r^j} \frac{\partial^{2(k-j)}}{\partial r^{2(k-j)}} \\
& - \sum_{k=0}^{2n+1} \frac{3^{2n+1} (2n+1)!}{k!(2n+1-k)!} \frac{1}{r^{2n+1-k}} \frac{\partial^{2n+1-k}}{\partial r^{2n+1-k}} \sum_{j=0}^k \frac{(-1)^{k-j} k!}{j!(k-j)!} \frac{\partial^{2j}}{\partial r^{2j}} \frac{L^{2(k-j)}}{r^{2(k-j)}} \\
& - \sum_{k=0}^{2n+1} \frac{3^{2n+1} (2n+1)!}{k!(2n+1-k)!} \frac{1}{r^{2n+1-k}} \frac{\partial^{2n+1-k}}{\partial r^{2n+1-k}} \sum_{j=0}^k \frac{(-1)^{k-j} k!}{j!(k-j)!} \frac{L^{2j}}{r^{2j}} \frac{\partial^{2(k-j)}}{\partial r^{2(k-j)}}
\end{aligned}$$

where $L^2 = -\frac{1}{2}(x_\mu \partial_\nu - x_\nu \partial_\mu)^2$. Clearly, also in the asymptotic limit, it seems technically difficult to find out a resolvent operator $\Delta_{R,y}$. On the other hand, in the deep IR limit, the expression (45) has a leading term that is nothing but a derivative of N-th order, where $N = 4n+2$. This implies that in the momentum space, $\Delta_{R,y}$ will trivially reduce to

$$\lim_{N \rightarrow \infty, p \rightarrow \infty} \Delta_{R,y} \rightarrow p^{-2} + O(p^{-N}) \quad (46)$$

($N > 3$), or

$$\lim_{p \rightarrow \infty} \Delta_{R,y} \rightarrow p^{-2} \sum_{N=0}^{\infty} b_N p^{-N} \quad (47)$$

(where b_N is an understood convolution of coefficients in (45) not important for our purposes). In other words, the classical configuration ϕ_0 has to satisfy the asymptotic Equation of motion

$$\sum_{m=0}^{\infty} A_m \square^{m+1} \phi_0 = M \delta(\zeta - z) \quad (48)$$

On the other hand, one can consider the UV asymptotic limit of our model, $r \rightarrow 0$.

In UV regime, $\square \rightarrow \infty$ so that the relevant lagrangian becomes

$$\mathcal{L} \rightarrow +\frac{1}{2} \varphi(\phi) \frac{\square}{e^{\frac{\square}{\Lambda_{NL}^2}}} \varphi(\phi) - V(\phi + \varphi(\phi)) \quad (49)$$

where the relevant functional φ is $\varphi(\phi) \rightarrow 3\lambda \frac{e^{\frac{\square}{\Lambda_{NL}^2}}}{\square} \phi^3 \dots$. This lagrangian can be explicitly rewritten as an expression of the following leading terms:

$$\begin{aligned}
\mathcal{L} \rightarrow & +\frac{9\lambda^2}{2} \frac{1}{n! \Lambda_{NL}^{2n}} \square^{n-1} \phi^6 - \frac{81\lambda^4}{2} \left(\frac{1}{n! \Lambda_{NL}^{2n}} \square^{n-1} \phi^3 \right)^4 \\
& - 27\lambda^4 \phi \left(\frac{1}{n! \Lambda_{NL}^{2n}} \square^{n-1} \phi^3 \right)^3 - \frac{54\lambda^2}{4} \phi^2 \left(\frac{1}{n! \Lambda_{NL}^{2n}} \square^{n-1} \phi^3 \right)^2 - \lambda \phi^3 \left(\frac{1}{n! \Lambda_{NL}^{2n}} \square^{n-1} \phi^3 \right) + \dots
\end{aligned} \quad (50)$$

As a consequence the asymptotic limit of $H_{R,y}$ for $r \rightarrow 0$, in spherical coordinates, has a form

$$\begin{aligned}
H_{R,y} \rightarrow & - \sum_k^N \alpha_{Nk} \frac{\partial^{2k}}{\partial r^{2k}} \sum_j^k \beta_{kj} \frac{3^j}{r^j} \frac{\partial^j}{\partial r^j} \frac{L^{2(k-j)}}{r^{2(k-j)}} \\
& - \sum_k^N \alpha_{Nk} \frac{\partial^{2k}}{\partial r^{2k}} \sum_j^k \beta_{kj} \frac{L^{2j}}{r^{2j}} \frac{3^{k-j}}{r^{k-j}} \frac{\partial^{k-j}}{\partial r^{k-j}} \\
& - \sum_k^N \alpha_{Nk} \frac{L^{2k}}{r^{2k}} \sum_j^k \beta_{kj} \frac{\partial^{2j}}{\partial r^{2j}} \frac{3^{k-j}}{r^{k-j}} \frac{\partial^{k-j}}{\partial r^{k-j}} \\
& - \sum_k^N \alpha_{Nk} \frac{L^{2k}}{r^{2k}} \sum_j^k \beta_{kj} \frac{3^j}{r^j} \frac{\partial^j}{\partial r^j} \frac{\partial^{2(k-j)}}{\partial r^{2(k-j)}} \\
& - \sum_k^N \alpha_{Nk} \frac{3^k}{r^k} \frac{\partial^k}{\partial r^k} \sum_j^k \beta_{kj} \frac{L^{2j}}{r^{2j}} \frac{\partial^{2(k-j)}}{\partial r^{2(k-j)}} \\
& - \sum_k^N \alpha_{Nk} \frac{3^k}{r^k} \frac{\partial^k}{\partial r^k} \sum_j^k \beta_{kj} \frac{\partial^{2j}}{\partial r^{2j}} \frac{L^{2(k-j)}}{r^{2(k-j)}}
\end{aligned} \tag{51}$$

where α_{Nk}, β_{kj} are combinations of factorials, not explicitly reported: they will be not important for our purposes. As a consequence, in deep UV regime, $H_{R,y}$ is a combination of all $\partial^M / \partial r^M$ and $1/r^P$ satisfying the constraint $M + P = 2N$:

$$\lim_{N \rightarrow \infty, r \rightarrow 0} H_{R,y} \sim - \lim_{N \rightarrow \infty} \sum_{M,P; M+P=2N} B_{MP} \frac{1}{r^P} \frac{\partial^M}{\partial r^M} \tag{52}$$

where B_{MP} is a combination of α, β coefficients. From this, we can see how divergent is the operator $H_{R,y}$ in the deep UV regime. As a consequence, we can find a trivial solution $\Delta_{R,y}$ for the problem Eq.(41) in the asymptotic limit $r \rightarrow 0$: just the trivial distribution $\Delta_{R,y} \rightarrow 0$.

Now, we can reconstruct a $\Delta_{R,y}$ connecting the two asymptotic solutions $r \rightarrow \infty$ and $r \rightarrow 0$ discussed above. The resolvent $\Delta_{R,y}$, in the momentum space, will have a form

$$\lim_{N \rightarrow \infty} \Delta_{R,y} \sim f(p) \mathcal{F}(p_* - p) \tag{53}$$

where $f(p)$ is a monotonically decreasing function with a limit $p \rightarrow \infty$ converging to a power series $p^{-2} \sum_{N=0} b_N p^{-N}$ of (47). $\mathcal{F}(p_* - p)$ is a distribution that has to satisfy the following asymptotic proprieties:

for $p < hp_*$

$$f(p) \mathcal{F}(p_* - p) = 0$$

for $p > kp_*$

$$f(p)\mathcal{F}(p_* - p) = qf(p)$$

where $q > 0$ is a constant, with k, h two real numbers satisfying $k \geq h > 0$. In the following discussion, we will assume $q = 1$: simply we can redefine $f(p)$ so that q is just embedded in this one. Let us note that, at priori, $f(p)$ can be a function defined in the range $(-\infty, p^*]$, *i.e* not necessary in $(-\infty, \infty)$ or $(-\infty, 0]$. These are all the general informations that we can get by our "deep" asymptotic limits. Among the possible solutions, one can propose as an *ansatz* that

$$\mathcal{F}(p_* - p) = \Theta(p^* - p)$$

where Θ is the Heaviside distribution. In fact, this case satisfies all conditions mentioned above. In particular, if $k = h = 1$ one obtain this result. Let us define $R = 1/p_*$. This definition is useful to get the regime in which surely one can consider a large class of $\mathcal{F}(p_* - p)$ as a Heaviside distribution $\Theta(p_* - p)$. In fact, $R/h < r < R/k$ is the region in which a generic \mathcal{F} deviates from an Heaviside step distribution: we can imagine it as a smoothed Heaviside (around the step). On the other hand, for $r \ll R/h$ and $r \gg R/k$, \mathcal{F} converges to a Heaviside distribution.

The fact that we have chosen the same labels R for the classicalon radius as well as for $R = 1/p_*$ is not a coincidence: the physical interpretation of R inside the distribution \mathcal{F} is exactly the classicalon radius. In fact, the transition of \mathcal{F} is related to the presence of a classical configuration of radius R . At this point, one can argue that the particular shape of \mathcal{F} around the Heaviside distribution has not a particular physical importance for our purposes: we are only interested to get qualitative proprieties of classicalization in scattering amplitudes. In a broad analogy, this is like the case of shape functions for atomic nuclei in which one can define an average radius R , with a monotonically decreasing behavior after R .

Let us discuss the important implications of these formal results on scattering amplitudes. In particular, we are interested to a $2 \rightarrow 2$ scattering as a significant example. So, let us consider the 4-correlator (4-Green's function):

$$\begin{aligned} \mathcal{G}(z_1, z_2, z_3, z_4) = & \frac{1}{Z} \int dR d^4y \mathcal{Y}(R, y) \frac{e^{-S[\phi_0(R, y)]}}{\sqrt{\text{Det}(H_{R, y})}} e^{-S_{int}} \\ & \times [\partial_\mu \Delta_{R, y}(z_1, z) \partial^\mu \Delta_{R, y}(z_2, z) \partial_\nu \Delta_{R, y}(z_3, z) \partial^\nu \Delta_{R, y}(z_4, z) + \text{permutations}] \end{aligned} \quad (54)$$

where Z is the partition function (40). At this point, we can find out some important qualitative proprieties of $\mathcal{G}(z_1, z_2, z_3, z_4)$ contained in our solution $\Delta_{R,y}$ of the form (53). Inserting the Heaviside-like distributions into (54), we will emerge a cutoff R inside the Green correlator. So that, from these very general and simple considerations, we can conclude that the maximal momentum probed by the $2 \rightarrow 2$ scattering can be just $p_* = 1/R$. As a consequence the 4-amplitude is

$$\lim_{p \rightarrow \infty, n \rightarrow \infty} \mathcal{M}_{\phi\phi \rightarrow \phi\phi} \leq a_n \frac{p_*^{2n}}{\Lambda_{NL}^{2n}} \quad (55)$$

We know the final result of our amplitude at 1-loop (31). Substituting $s = p_*^2 \leq \Lambda_{NL}^2$ in the series, the convergence can be manifestly shown. Infact, $a_n < 1/(n+1)!n$, with a_n the coefficient of our series: for the direct comparison test, convergence of $\sum_n 1/(n+1)!n$ implies convergence of $\sum_n a_n$. For example, in s-channel, we have

$$\lim_{s \rightarrow \Lambda^2} \mathcal{A} \rightarrow \frac{9\lambda^2}{4\pi^2} \sum_{n=0}^{\infty} \frac{1 - 2^{-n}}{(n+1)!n} (\tilde{s}^*)^n \quad (56)$$

and the following bound:

$$\lim_{s > \Lambda^2} \mathcal{A} \leq \frac{9\lambda^2}{4\pi^2} \sum_{n=0}^{\infty} \frac{1 - 2^{-n}}{(n+1)!n} (\tilde{s}^*)^n \quad (57)$$

where \tilde{s} is the adimensional Mandelstam variable $\tilde{s} = s/\Lambda_{NL}^2$, so that $\tilde{s}^* = s/\Lambda_{NL}^2 = (p^*)^2/\Lambda_{NL}^2$ is the cut-off value in the s-channel. The convergence of this amplitude to a finite one is manifest for $\tilde{s}^* \leq 1$. Analogous considerations are valid for t- and u-channels. More surprisingly, the amplitude (31) can converge also for $\tilde{s}^* = \text{const} > 1$, thanks to the factor $1/(n+1)!$.

3.3 Further comments and implications

In this section, we would like to briefly comment about some implications of classicalization in non-local models. These aspects will deserve deeper analysis and future investigations beyond the purposes of this paper.

1) Our result can be extended for vector-bosons' scatterings. As shown in [4, 5, 6, 7, 8], a non-local gauge theory can be formulated. Scattering like $VV \rightarrow VV$, where V is a vector boson, for $E \gg \Lambda_{NL}$ behaves similarly to the one considered here: an infinite series of divergences will emerge [9, 10, 23].

2) UV fate of non-local models, *i.e* how to decide if a non-local theory is Wilson-like UV completed or Dvali-like one. We can argue that this problem is not different in our

case with respect to local QFTs. These aspects were just discussed in various papers cited above. Because of this, we will not discuss this issue in our paper.

3) Classicalization can help non-local models to eliminate acausal divergences in scatterings, *i.e* it is an unitarization and causalization of scatterings. However, divergences coming from radiative corrections of propagators or vacuum polarization diagrams seem to remain still alive! These kinds of divergences were considered in our paper [23] (in contest of $\mathcal{N} = 1$ susy non-local models). These divergences can have observable effects in the running coupling constants. Clearly, these are suppressed as a powers Λ_{NL}^{-n} so that these effects are negligible for $E \ll \Lambda_{NL}$. $\mathcal{N} = 1$ (rigid) supersymmetry seems to help to cancel an infinite number of radiative divergences, as explicitly shown in [23], even if not all the infinite ones! (Curiously) $\mathcal{N} > 1$ susy non-local models were not studied in literature. One can conjecture that for $\mathcal{N} = 2$ susy non-local models, more divergences can be eliminated.

4) EWMKEJ model is manifestly Lorentz invariant and CPT invariant at three level. However, one can argue that classicalization can be a more general procedure for a more general class of non-local models without Lorentz and CPT invariance. This can be strongly motivated by QFT in non-commutative geometries in which non-local terms in matter generically emerge ^{4 5}.

5) In our considerations, we have not considered gravity. However, one can argue that these results can be also extended for non-local modifications of General-Relativity [28, 29, 30, 31, 32, 33, 34, 35, 36, 37, 38, 39, 40, 41, 42]⁶. In these model, GR singularities are removed, but quantization is still inconsistent. However, classicalons' formation in graviton-graviton scatterings can unitarize a non-local extension of GR. As notice by Dvali and collaborators, a classicalon can be nothing but a black hole in gravitational scatterings [50]. A Black hole can be formed in graviton-graviton scatterings, unitarizing a gravitational theory.

6) In the case of $\Lambda_{NL} \simeq 100$ TeV, one can speculate about implications for UHECR or for future colliders. In particular: i) classicalon resonance can give characteristic signatures in collisions, as proposed in papers about classicalizations cited above; ii)

⁴For a recent new model of non-commutative QFT see [46].

⁵An intriguing area to explore could be non-local models in which Lorentz Invariance is an emergent symmetry. See [47, 48, 49] for examples of quantum field theories in which Lorentz invariance is not a fundamental symmetry.

⁶We would like to mention that recently a study of external geodetic stability in particular branches of massive gravity [43]. In subregions of parameters of these models, naked singularities can exist. This can be connected to the existence of new items called *frizzyballs* [44, 45].

as mentioned above, polynomial corrections to the cross sections can be a clear signal beyond a local quantum field theory [9, 10, 25, 26, 27].

4 Conclusion and remarks

In this paper, we have shown how classicalization can cure acausal divergences of non-local QFT coming in "trans-non-local" limit. In particular, we have discussed a particular class of non-local QFT for a scalar field well studied in literature cited above. We have explicitly shown how the formation of a classicalon can avoid acausalities in scatterings for an energy higher than the non-locality scale. We have also discussed possible implications for gauge theories, gravity, cosmology, UHECR and future 100 TeV-colliders. We conclude that classicalization seems a natural UV completion of non-local quantum field theories, unitarizing and causalizing their scattering amplitudes.

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References

- [1] G. Dvali, D. Pirtskhalava, Phys.Lett. B699 (2011) 78-86, arXiv:1011.0114 [hep-ph];
G. Dvali, arXiv:1101.2661 [hep-th];
G. Dvali, A. Franca and C. Gomez, arXiv:1204.6388 [hep-th];
G. Dvali and C. Gomez, JCAP **1207** (2012) 015 [arXiv:1205.2540 [hep-ph]].
- [2] B. Bajc, A. Momen and G. Senjanovic, arXiv:1102.3679 [hep-ph].
- [3] C. Grojean and R. S. Gupta, JHEP **1205** (2012) 114 [arXiv:1110.5317 [hep-ph]];
J. Rizos and N. Tetradis, JHEP **1204** (2012) 110 [arXiv:1112.5546 [hep-th]];
N. Brouzakis, J. Rizos, N. Tetradis, Phys.Lett. B708 (2012) 170-173, arXiv:1109.6174 [hep-th];
R. Akhoury, S. Mukohyama and R. Saotome, arXiv:1109.3820 [hep-th];

- R. Percacci and L. Rachwal, Phys. Lett. B **711** (2012) 184 [arXiv:1202.1101 [hep-th]];
- P. Asimakis, N. Brouzakis, A. Katsis and N. Tetradis, Phys. Lett. B **743** (2015) 75 [arXiv:1412.4275 [hep-th]];
- R. Percacci and L. Rachwal, PoS CORFU **2011** (2011) 100;
- A. Vikman, Europhys. Lett. **101** (2013) 34001 [arXiv:1208.3647 [hep-th]];
- A. Kovner and M. Lublinsky, JHEP **1211** (2012) 030 [arXiv:1207.5037 [hep-th]];
- J. Rizos, N. Tetradis and G. Tsolias, JHEP **1208** (2012) 054 [arXiv:1206.3785 [hep-th]].
- [4] D. A. Eliezer and R. P. Woodard, Phys. Rev. D **40**, 465 (1989).
- [5] D. A. Eliezer and R. P. Woodard, Nucl. Phys. B **325**, 389 (1989).
- [6] J. W. Moffat, Phys. Rev. D **41**, 1177 (1990).
- [7] G. Kleppe and R. P. Woodard, Nucl. Phys. B **388**, 81 (1992).
- [8] D. Evens, J. W. Moffat, G. Kleppe, and R. P. Woodard, Phys. Rev. D **43**, 499 (1991).
- [9] A. Jain and S. D. Joglekar, Int. J. Mod. Phys. A **19**, 3409 (2004).
- [10] S. D. Joglekar, hep-th/0601006.
- [11] F. Brischese, A. Marcian, L. Modesto and E. N. Saridakis, Phys. Rev. D **87** (2013) 8, 083507 [arXiv:1212.3611 [hep-th]].
- [12] L. Modesto, arXiv:1302.6348 [hep-th].
- [13] C. Bambi, D. Malafarina and L. Modesto, Phys. Rev. D **88** (2013) 044009 [arXiv:1305.4790 [gr-qc]].
- [14] L. Modesto, arXiv:1305.6741 [hep-th].
- [15] C. Bambi, D. Malafarina, A. Marcian and L. Modesto, Phys. Lett. B **734** (2014) 27 [arXiv:1402.5719 [gr-qc]].
- [16] L. Modesto, arXiv:1402.6795 [hep-th].
- [17] L. Modesto and L. Rachwal, Nucl. Phys. B **889** (2014) 228 [arXiv:1407.8036 [hep-th]].

- [18] G. Calcagni and L. Modesto, Phys. Rev. D **91** (2015) 12, 124059 [arXiv:1404.2137 [hep-th]].
- [19] L. Modesto and L. Rachwal, arXiv:1503.00261 [hep-th].
- [20] P. Don, S. Giaccari, L. Modesto, L. Rachwal and Y. Zhu, JHEP **1508** (2015) 038 [arXiv:1506.04589 [hep-th]].
- [21] N.N.Bogolibov and D.V.Shirkov, ‘Introduction to Theory of Quantized Fields’ (3rd ed.), John Wiley(1980). See pg. 200-220.
- [22] S. Weinberg, The Quantum Theory of Fields, Volume 1: Foundations (Cambridge University press, Cambridge, 2005) Chapter 4.
- [23] A. Addazi and G. Esposito, Int. J. Mod. Phys. A **30** (2015) 1550103 [arXiv:1502.01471 [hep-th]].
- [24] H. J. Wannig, J. Mod. Phys. **6** (2015) 670 [arXiv:1301.0258 [hep-ph]].
- [25] S. D. Joglekar, Int. J. Theor. Phys. **47** (2008) 2824 [arXiv:0704.0995 [hep-ph]].
- [26] A. Haque and S. D. Joglekar, J. Phys. A **41** (2008) 215402 [hep-th/0701171].
- [27] T. Biswas, E. Gerwick, T. Koivisto and A. Mazumdar, Phys. Rev. Lett. **108**, 031101 (2012);
T. Biswas, A. Mazumdar and W. Siegel, JCAP **0603**, 009 (2006);
T. Biswas and N. Okada, arXiv:1407.3331 [hep-ph].
- [28] N. Arkani-Hamed, S. Dimopoulos, G. Dvali and G. Gabadadze, hep-th/0209227.
- [29] C. Deffayet and R. P. Woodard, JCAP **0908** (2009) 023 [arXiv:0904.0961 [gr-qc]].
- [30] T. Koivisto, Phys. Rev. D **77** (2008) 123513 [arXiv:0803.3399 [gr-qc]].
- [31] T. S. Koivisto, Phys. Rev. D **78** (2008) 123505 [arXiv:0807.3778 [gr-qc]].
- [32] S. Jhingan, S. Nojiri, S. D. Odintsov, M. Sami, I. Thongkool and S. Zerbini, Phys. Lett. B **663** (2008) 424 [arXiv:0803.2613 [hep-th]].
- [33] K. Bamba, S. Nojiri, S. D. Odintsov and M. Sasaki, Gen. Rel. Grav. **44** (2012) 1321 [arXiv:1104.2692 [hep-th]].
- [34] Phys. Lett. B **659** (2008) 821 [arXiv:0708.0924 [hep-th]].

- [35] S. Capozziello, E. Elizalde, S. Nojiri and S. D. Odintsov, Phys. Lett. B **671** (2009) 193 [arXiv:0809.1535 [hep-th]].
- [36] N. A. Koshelev, Grav. Cosmol. **15** (2009) 220 [arXiv:0809.4927 [gr-qc]].
- [37] R. P. Woodard, Found. Phys. **44** (2014) 213 [arXiv:1401.0254 [astro-ph.CO]].
- [38] G. Cusin, J. Fumagalli and M. Maggiore, JHEP **1409** (2014) 181 [arXiv:1407.5580 [hep-th]].
- [39] Y. Dirian and E. Mitsou, JCAP **1410** (2014) 10, 065 [arXiv:1408.5058 [gr-qc]].
- [40] S. Talaganis, T. Biswas and A. Mazumdar, arXiv:1412.3467 [hep-th].
- [41] A. Conroy, T. Koivisto, A. Mazumdar and A. Teimouri, Class. Quant. Grav. **32** (2015) 1, 015024 [arXiv:1406.4998 [hep-th]].
- [42] D. Chialva and A. Mazumdar, Mod. Phys. Lett. A **30** (2015) 03n04, 1540008 [arXiv:1405.0513 [hep-th]].
- [43] A. Addazi and S. Capozziello, Int. J. Theor. Phys. **54** (2015) 6, 1818 [arXiv:1407.4840 [gr-qc]].
- [44] A. Addazi, arXiv:1508.04054 [gr-qc].
- [45] A. Addazi, arXiv:1510.05876 [gr-qc].
- [46] S. Kawamoto and T. Kuroki, arXiv:1503.08411 [hep-th].
- [47] W. Heisenberg, Rev. Mod. Phys. **29** (1957) 269;
J. Bjorken, Ann. Phys. **24**, 174 (1963);
T. Eguchi, Phys. Rev. D **14** (1976) 2755.
- [48] C. Froggatt and H. B. Nielsen, *Origin of Symmetries* (World Scientific, 1991);
S. Chadha and H. B. Nielsen, Nucl. Phys. **B217**, 125 (1983);
J. Bjorken, *Proceedings of the Les Arcs Conference on New and Exotic Phenomena*, (Editions Frontieres, 1987), p. 1.
- [49] J.L. Chkareuli, C.D. Froggatt, H.B. Nielsen, Phys. Rev. Lett. **87** (2001) 091601;
J.L. Chkareuli, C.D. Froggatt, H.B. Nielsen, Nucl. Phys. B **609** (2001) 46;
P. Kraus, E.T. Tomboulis, Phys. Rev. D **66** (2002) 045015;

- A. Jenkins, Phys. Rev. D **69** (2004) 105007;
 J. L. Chkareuli, C. D. Froggatt, J. G. Jejelava and H. B. Nielsen, Nucl. Phys. B **796** (2008) 211 [arXiv:0710.3479 [hep-th]];
 J.L. Chkareuli, C.D. Froggatt, R.N. Mohapatra, H.B. Nielsen, hep-th/0412225;
 A.T. Azatov, J.L. Chkareuli, Phys. Rev. D **73** (2006) 065026;
 J. L. Chkareuli, arXiv:1206.1368 [hep-th];
 J. L. Chkareuli and Z. Kepuladze, Eur. Phys. J. C **72** (2012) 1954 [arXiv:1108.0399 [hep-ph]];
 J. L. Chkareuli, arXiv:1410.6837 [hep-th].
- [50] G. Dvali and C. Gomez, arXiv:1005.3497 [hep-th];
 G. Dvali, S. Folkerts and C. Germani, Phys. Rev. D **84** (2011) 024039 [arXiv:1006.0984 [hep-th]];
 E. Spallucci and S. Ansoldi, Phys. Lett. B **701** (2011) 471 [arXiv:1101.2760 [hep-th]];
 F. Berkhahn, D. D. Dietrich, S. Hofmann, Phys.Rev.Lett. **106** (2011) 191102, arXiv:1102.0313 [hep-th];
 F. Berkhahn, D. D. Dietrich, S. Hofmann JCAP **1109** (2011) 024 e-Print: arXiv:1104.2534 [hep-th];
 F. Berkhahn, S. Hofmann, F. Kuhnel, P. Moyassari and D. Dietrich, Phys. Rev. Lett. **108** (2012) 131102 [arXiv:1106.3566 [hep-th]].
- [51] A. Addazi and M. Bianchi, JHEP **1412** (2014) 089 [arXiv:1407.2897 [hep-ph]].
- [52] A. Addazi, JHEP **1504** (2015) 153 [arXiv:1501.04660 [hep-ph]].
- [53] A. Addazi and M. Bianchi, JHEP **1507** (2015) 144 [arXiv:1502.01531 [hep-ph]].
- [54] A. Addazi and M. Bianchi, JHEP **1506** (2015) 012 [arXiv:1502.08041 [hep-ph]].
- [55] A. Addazi, arXiv:1504.06799 [hep-ph].
- [56] A. Addazi, arXiv:1505.00625 [hep-ph].
- [57] A. Addazi, arXiv:1505.02080 [hep-ph].
- [58] A. Addazi, arXiv:1506.06351 [hep-ph].
- [59] A. Addazi, M. Bianchi and G. Ricciardi, arXiv:1510.00243 [hep-ph].

[60] A. Addazi, arXiv:1510.02911 [hep-ph].